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Theory of magnetoelectric photocurrent generated by direct interband transitions in a semiconductor quantum well

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A linearly polarized light normally incident on a semiconductor quantum well with spin-orbit coupling may generate pure spin current via direct interband optical transition. An electric photocurrent can be extracted from the pure spin current when an in-plane magnetic field is applied, which has been recently observed in the InGaAs/InAlAs quantum well [Dai et al., Phys. Rev. Lett. 104, 246601 (2010)]. Here we present a theoretical study of this magnetoelectric photocurrent effect associated with the interband transition. By employing the density matrix formalism, we show that the photoexcited carrier density has an anisotropic distribution in k space, strongly dependent on the orientation of the electron wavevector and the polarization of the light. This anisotropy provides an intuitive picture of the observed dependence of the photocurrent on the magnetic field and the polarization of the light. We also show that the ratio of the pure spin photocurrent to the magnetoelectric photocurrent is approximately equal to the ratio of the kinetic energy to the Zeeman energy, which enables us to estimate the magnitude of the pure spin photocurrent. The photocurrent density calculated with the help of an anisotropic Rashba model and the Kohn-Luttinger model can produce all three terms in the fitting formula for measured current, with comparable order of magnitude, but discrepancies are still present and further investigation is needed.

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I. INTRODUCTION

Optical injection and detection of spin current have attracted a lot of interest recently in the context of nonmagnetic semiconductor spintronics.1,2 Exciting progress has been made in recent years, such as the injection of spin-polarized currents by using the spin galvanic effect1 and the circular photogalvanic effect,3 the observation of spin accumulations induced by the spin Hall effect with the help of the Kerr rotation5 and the p-n junction light-emitting diode,6 the inverse spin Hall effect,7,8 and the detection of pure spin photocurrent by the second-harmonic effect.9,10

Generation and detection of the pure spin current where there is no net charge current is one of the challenging issues in spintronics. Optically the generation of pure spin photocurrent has been realized by injecting orthogonally polarized one- and two-photon linear lights,11–15 or by the incidence of unpolarized or linearly polarized lights into bulk III–V semiconductors and quantum wells with spin-orbit couplings.16–21

In the latter case, a linearly polarized light can be decomposed into circularly polarized lights with opposite helicities. Due to the conservation of angular momentum, the circularly polarized lights with opposite helicities excite carriers with opposite spins, which are locked to opposite momenta owing to the spin-orbit coupling. As a result, the photo-injected carriers with opposite spins always have opposite velocities, leading to the pure spin photocurrent. One way to observe this underlying pure spin photocurrent is to apply an in-plane magnetic field, which will create an imbalance of spins. The spin imbalance leads to the asymmetry of electron velocities by the spin-orbit coupling and can extract an electric photocurrent from the pure spin photocurrent. The field-induced conversion from pure spin photocurrent to electric photocurrent (also referred to as the magneto-gyrotropic photogalvanic effect [MPGE]) was systematically studied via the intraband and intersubband electron heating by THz and microwave radiations, in which spin-dependent scatterings are necessary for the observed charge current.

Very recently the magnetoelectric photocurrent has been observed via direct interband optical transitions.22–25 Unlike the intraband transition, in the direct interband transition the electron momentum is conserved, and the magnetoelectric photocurrent is a few orders of magnitude larger than that via the intraband transition. Therefore, it appears to be promising as a simple and practical method to generate and detect the spin current. Although symmetry analysis23 can justify the observed electric photocurrent as a function of magnetic fields and polarization of the light, a microscopic description is still needed. In particular, the current induced by a parallel field is against intuition, considering the form of the spin-orbit coupling derived from the symmetry argument. Moreover, how to estimate the zero-field pure spin photocurrent from the magnetoelectric photocurrent is an interesting issue.

In this work we study theoretically the generation of photocurrent via direct interband transitions excited by shedding a linearly polarized light normally into a semiconductor quantum well. We find that the k space anisotropy of the carrier density provides an intuitive microscopic picture for the magnetoelectric photocurrent and the zero-field pure spin photocurrent. The origin of the observed dependences of the photocurrent on the magnetic field and the polarization of the light can be illustrated within this picture. The photocurrent density calculated with the help of a minimal model with an anisotropic Rashba model and the Kohn-Luttinger model is comparable with the experiment. Moreover, we propose an approach to estimate the magnitude of the undetectable pure

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spin photocurrent, which is generated at zero magnetic field by the same linear light, from the observed magnetoelectric photocurrent. Part of the results was very briefly reported in Ref. 25.

The paper is organized as follows. In Sec. II, we review key features of the experimentally observed magnetoelectric photocurrent as a function of magnetic fields and the polarization of the linear light and present a symmetry argument. In Sec. III, a minimal model of the quantum well is presented, as well as the electric-radiation interaction that accounts for the interband optical transitions. In Sec. IV, we introduce the anisotropic photoexcited carrier density in \(\mathbf{k}\) space, with the help of a standard density matrix formalism. In Sec. V A, we illustrate the origin of the field and polarization dependence of the magnetoelectric photocurrent in terms of the anisotropic photoexcited carrier density. In Sec. V B, we compare the calculated magnetoelectric photocurrent densities with the experiment. In Sec. VI, the estimate of the underlying zero-field pure spin photocurrent is discussed. Finally, a summary is given in Sec. VII.

II. EXPERIMENT AND SYMMETRY ANALYSIS

In this section we will review the experiment on the magnetoelectric photocurrent generated via direct interband optical transitions by shedding a linearly polarized light normally incident into an InGaAs/InAlAs quantum well.25 The experiment setup reported in Ref. 25 is schematically illustrated in Fig. 1.

In the experiment, the current density measured along the \(x\) direction can be formulated as

\[
j_x = c_0 B_y + c_x B_x \cos 2\theta + c_y B_y \sin 2\theta, \tag{1}\]

where \(B_x\) and \(B_y\) are the magnetic fields along the \(x\) and \(y\) directions, respectively. \(\theta\) is the polarization angle of the linearly polarized light with respect to the \(x\) axis. \(c_{0/x/y}\) are constant coefficients linearly scaled with the light power. The \(x\) and \(y\) axes here are defined along \([1\bar{1}0]\) and \([110]\) crystallographic directions of the zincblende structure, respectively.

The experimental results are consistent with the symmetry argument of the \(C_{2v}\) point group.23 Inversion-asymmetric zincblende heterostructures grown along the \([001]\) crystallographic direction usually have the \(C_{2v}\) point group symmetry. As a result, the electric photocurrent induced by in-plane magnetic fields can be phenomenologically written as (refer to Appendix A for details)

\[
j_\mu = \sum_v (\chi^{\mu\nu\bar{\nu}} B_\mu E_\nu + \chi^{\mu\bar{\nu}\nu} B_\bar{\nu} E_\nu), \tag{2}\]

where \(\mu, \nu\) run over \(\{x, y\}\), and \(\bar{\mu}, \bar{\nu} = y\) if \(\mu, \nu = x\), and vice versa. \(\chi\) is a fourth-rank pseudotensor that relates the electric photocurrent to the polarization electric field components \((E_x, E_y) \propto (\cos \theta, \sin \theta)\) of the incident light and the applied magnetic fields \((B_x, B_y)\). One can readily check that Eq. (2) yields the same form as Eq. (1).

Another consequence of the \(C_{2v}\) symmetry is that when \(x||[110]\) and \(y||[1\bar{1}0]\), the invariants of the \(C_{2v}\) symmetry allow the spin-orbit coupling in linear \(\mathbf{k}\) to be only in the form26,27 (refer to Appendix A for details)

\[
H_{SOC} = \lambda_{x,y} k_x - \lambda_{x,y} k_y, \tag{3}\]

where \(k_{x,y}\) is the wavevector along the \(x\) (\(y\)) axis, the Pauli matrix \(\sigma_{x,y}\) depicts the spin along the \(x\) (\(y\)) direction, and \(\lambda_{x,y}\) is the spin-orbit coupling coefficient along the \(x\) (\(y\)) direction. Usually, the current along \(x\) direction corresponds to the average shift of momenta of carriers along the \(x\) direction. According to Eq. (3), \(k_x\) couples only to the spin along \(y\) directions, thus it can be shifted only by a magnetic field along the \(y\) direction. A magnetic field applied along the \(x\) direction cannot shift the average \(k_x\) [see Fig. 2(b)]. According to this argument, the last term of Eq. (1) seems against intuition because it originates from a magnetic field along the \(x\) direction.

The main task of this work is to present an intuitive picture to solve the above dilemma, as well as the magnetic field and polarization angle dependence of the electric photocurrent, in terms of the anisotropy of the photoexcited carrier density in \(\mathbf{k}\) space. We will see that the three terms in Eq. (1) can be well interpreted by the microscopic picture presented in Fig. 4 (subfigures in boxes).

III. MODEL HAMILTONIAN

In this work we consider a minimal model that describes the energy bands near the band gap \(\Delta\) of the quantum well. The advantage of the minimal model is that it is analytically solvable, thus physically transparent in delivering the physical picture. We expect that more sophisticated calculations, e.g., by using the 14-band \(\mathbf{k} \cdot \mathbf{p}\) model or the full band structure local density approximation,28 can cover more details but will not deviate from the physical picture obtained by the minimal model.

The total Hamiltonian is given by

\[
H = H_C + H_V + \tilde{V}(t), \tag{4}\]

where \(H_C\) is for the conduction bands, \(H_V\) for the valence bands, and \(\tilde{V}(t)\) for the electric-radiation that couples to the average shift of momenta of carriers along the \(x\) direction.
the conduction and valence bands by the linearly polarized light. The lowest conduction subbands are described by a two-dimensional free-electron gas with the Rashba and Dresselhaus spin-orbit couplings and in-plane magnetic fields,

\[ H_C = \frac{\hbar^2}{2m^*} (k_x^2 + k_y^2) + (\lambda_x, \sigma_x, k_x - \lambda_s, \sigma_y, k_y) + h \cdot \sigma, \]

where \( m^* \) is the effective mass of electron, \( \hbar \) is Planck’s constant over \( 2\pi \), \( k_x \) is the wavevector along the \( x \) (\( y \)) axis, \( \lambda_x = \alpha + \beta \), \( \lambda_s = \alpha - \beta \), and \( \alpha \) and \( \beta \) are the Rashba and Dresselhaus spin-orbit coupling coefficients, respectively. Throughout the work, \( x \) and \( y \) axes are defined along \([110] \) and \([100] \) crystallographic directions, respectively. \( \sigma = (\sigma_x, \sigma_y) \) is the vector of the Pauli matrices, \( h = (h_x, h_y) = \frac{1}{2} g_s \mu_B (B_x, B_y) \) is the Zeeman energies induced by the in-plane magnetic fields \( (B_x, B_y) \), with \( g_s \) the Landé\’s factor, and \( \mu_B \) the Bohr magneton. The eigenenergies and eigenstates of \( H_C \) are given by

\[ \epsilon_{\pm}(k) = \frac{\hbar^2}{2m^*} k^2 \pm \Lambda, \]

\[ |+\rangle = \left[ \begin{array}{c} \sqrt{2} \gamma_1 \\ \gamma_2 \end{array} \right], \quad |−\rangle = \left[ \begin{array}{c} \gamma_1 \\ \sqrt{2} \gamma_2 \end{array} \right], \]

where \( \gamma_1 \) and \( \gamma_2 \) are two model parameters, and \( k = (k_x, k_y, k_z) \) is the wavevector. For simplicity, we approximate \( (k_z) = 0 \) and \( (k_x^2) \gg (k_y^2) \) in \( H_V \) when considering the quantization along the \( z \) direction of the quantum well, where \( d \) is the thickness of the quantum well. The eigenenergies and states of the valence bands are given by

\[ \epsilon_{1/2}(k) = \epsilon_{3/2}(k) = -\Delta - \frac{\hbar^2 \gamma_1}{2m^*} (k_z^2 + (k_x^2)) \pm \frac{\hbar^2 \gamma_2}{2m^*} \Omega, \]

\[ |1, k\rangle = \left( \begin{array}{c} Q \\ 0 \\ K \end{array} \right), \quad |2, k\rangle = \left( \begin{array}{c} 0 \\ 0 \\ -K \end{array} \right), \]

\[ |3, k\rangle = \left( \begin{array}{c} -K^* \\ 0 \\ Q \end{array} \right), \quad |4, k\rangle = \left( \begin{array}{c} 0 \\ K^* \\ Q \end{array} \right), \]

where \( \Delta \) is the energy gap, and \( Q = C_e(-k^2 + 2(k_x^2) + \Omega), \ K = \sqrt{3} C_e k_x^2, \ \Omega = 2\sqrt{2} k^4 + (k_x^2)^2 - k_x^2 (k_y^2), \) and \( C_e = |3k^4 + (k^2)^2 - (k_x^2)^2 - (k_y^2)^2|^{1/2}. \)

The electric-radiation interaction that couples the conduction and valence bands is given by\(^{29}\)

\[ \hat{V}(t) = \frac{e}{m_e} \hat{A} \cdot \hat{p}, \]

where \( e \) is the electron charge and \( \hat{p} \) the electron momentum operator. Under the electric dipole approximation, the vector potential \( \hat{A} \) is related to the electric field of the single-frequency light by

\[ \hat{A}(t) = \frac{1}{i\omega} [E(\omega)e^{-i\omega t} - E^*(\omega)e^{i\omega t}], \]

where we consider a linearly polarized single-color light incident normally to the \( x-y \) plane, with the frequency \( \omega \) and the polarization electric vector

\[ E = E_0(\cos \theta, \sin \theta, 0), \]

where \( E_0 \) is the amplitude and \( \theta \) is the polarization angle with respect to the \( x \) axis. With the help of the Poynting vector, the electric component \( E_0 \) can be evaluated from the energy flux of the laser by \( \xi I = (1/2)\sqrt{\varepsilon_0/\mu_0} E_0^2 \), where \( \xi \) is the absorption efficiency and \( I \) is the light intensity. \( \varepsilon_0 \) and \( \mu_0 \) are the dielectric constant and magnetic permeability in the vacuum, respectively. We always assume \( \omega > \Delta \), therefore the dominant optical absorption mechanism is the direct interband transition. “Direct” means that the wavevector \( k \) of the electron keeps unchanged in the transition.

We neglect the Zeeman effect in the valence bands because the contribution to currents from holes is expected to be much smaller than electrons considering the short charge and spin lifetimes of holes in \( n \)-type quantum wells. Besides the lowest subbands \( |+\rangle \) and \( |-\rangle \), another pair of conduction subbands \( |+\rangle \) and \( |-\rangle \) above them are also considered in the numerical calculation. Approximated as the energy levels of an infinite depth potential well, \( |\pm\rangle \) are about \( 3(h^2/2m^*) (\sigma/d)^2 \) above \( |\pm\rangle \). The differences in the effective mass and spin-orbit couplings between \(|\pm\rangle \) and \(|\pm\rangle \) are neglected for simplicity. Due to the parity in the \( z \) direction, conduction bands \(|+\rangle \) and \(|-\rangle \) couple to valence bands \(|1\rangle \) and \(|4\rangle \), while \(|+\rangle \) and \(|-\rangle \) to \(|2\rangle \) \((k_x \) \) and \(|3\rangle \). For simplicity, we also neglect the diamagnetic contribution,\(^{30-32}\) which should give qualitatively similar results to the Zeeman effect.

**IV. ANISOTROPY OF PHOTOEXCITED CARRIER DENSITY IN \( k \) SPACE**

The photoexcited carrier density and all the physical quantities can be found within the density matrix formalism. In this work we will consider only the steady-state nonequilibrium photoexcited carrier density, which can be found by the approach similar to that of the second-order nonlinear optical susceptibilities.\(^{33}\) We start with the Liouville equation of the density matrix,

\[ \partial_t \rho_{nm} = -i \hbar [H, \rho]_{nm} - \gamma (\rho_{nm} - \rho_{nm}^{(0)}), \]

where \( n \) and \( m \) run over the states in Eqs. (7) and (10), and \( \gamma \) is a phenomenological damping parameter. \( \rho_{nm}^{(0)} \) is the equilibrium density matrix before the light excitation. Because we assume a \( n \)-type quantum well, the valence bands are
fully occupied $\rho_{v,v'}^{(0)} = \delta_{v,v'}$ where $v,v' \in \{1,2,3,4\}$. While the initial equilibrium density matrix of the conduction bands are described by the Fermi function,

$$\rho_{v,v}^{(0)}(k) = f(e_{v}(k))\delta_{v,v'} \equiv \frac{\delta_{v,v'}}{e_{v}(k) - E_F(k) + \frac{i}{\hbar}T} + 1,$$

(15)

where $c,c' \in \{+, -, +', -'\}$, $E_F$ is the Fermi energy, $k_B$ is the Boltzmann constant, and $T$ is the temperature. By treating $|c,k\rangle$ and $|v,k\rangle$ as unperturbed part and $V(t)$ as perturbation, the perturbation equations up to the second order are given by

$$\begin{align*}
\partial_t \rho_{nm}^{(0)} &= -i \omega_{nm} \rho_{nm}^{(0)}, \\
\partial_t \rho_{nm}^{(1)} &= -i \omega_{nm} \rho_{nm}^{(1)} - \frac{i}{\hbar} \left[ \hat{V}, \rho_{nm}^{(0)} \right]_{\text{ann}} - \gamma \rho_{nm}^{(1)}, \\
\partial_t \rho_{nm}^{(2)} &= -i \omega_{nm} \rho_{nm}^{(2)} - \frac{i}{\hbar} \left[ \hat{V}, \rho_{nm}^{(1)} \right]_{\text{ann}} - \gamma \rho_{nm}^{(2)},
\end{align*}$$

(16)

where $\omega_{nm}(k) = [\epsilon_n(k) - \epsilon_m(k)]/\hbar$. $\rho_{nm}$ are functions of $k$ because the light is momentum-free under the electric dipole approximation. After a straightforward derivation, the leading order of the light-induced steady-state density matrix of the conduction bands is found to be of the second order:

$$\rho_{c,c}^{(2)}(k) = \frac{\tau_c \pi \epsilon_c^2}{\hbar^2 \omega_c^2} \sum_v \left[ \rho_{c,c}^{(2)}(k) \cdot \mathbf{E}(\omega) \right] \rho_{c,c}(k) \cdot \mathbf{E}^*(\omega),$$

(17)

where the steady state is approximated by introducing the momentum relaxation time $\tau_c$.

$$\rho_{c,c}^{(2)}(k) \approx \tau_c \partial_t \rho_{c,c}^{(2)}(k,t).$$

(18)

$\tau_c = \mu_m m^*/e$ can be estimated from the mobility $\mu_m$ and the effective mass $m^*$. Equation (17) recovers the result by employing the semiconductor optical Bloch equations and Fermi’s golden rules. In the following, we will suppress the superscript of $\rho_{c,c}^{(2)}$ for simplicity. By substituting the velocity $v_{c,c}$ in Eq. (17) by the position $r_{c,c}$ according to

$$\langle c,k|\hat{r}_c|v,k \rangle = \langle c,k|\frac{1}{i\hbar}[\hat{r}_c,H_0]|v,k \rangle = i \omega_{c,c} \langle c,k|\hat{r}_c|v,k \rangle,$$

(19)

Eq. (17) can also be expressed as

$$\rho_{c,c}^{(2)}(k) = \frac{\tau_c \pi \epsilon_c^2}{\hbar^2 \omega_c} \sum_v \left[ \omega_{c,c}(1 - f_c)\delta(\omega - \omega_{c,c}) \\
+ \omega_{c,c}(1 - f_c)\delta(\omega - \omega_{c,c}) \right]_{\text{ann}} \left[ r_{c,c}(k) \cdot \mathbf{E}(\omega) \right],$$

(20)

By using the eigenstates in Eqs. (7) and (10) and considering the spatial wavefunctions of the eigenstates, the elements of the transition matrix $r_{c,c} \equiv (x_{c,c},y_{c,c}) = (x_{c,c},y_{c,c})$ are found as

$$\begin{align*}
x_{c,c} &= \langle c,k|\xi|c,k \rangle, \\
y_{c,c} &= \langle c,k|\eta|c,k \rangle.
\end{align*}$$

(21)

and

$$\begin{align*}
x_{c,c} &= \langle c,k|\xi|c,k \rangle, \\
y_{c,c} &= \langle c,k|\eta|c,k \rangle.
\end{align*}$$

(22)

where we have defined the effective dipole length

$$a_{cv} \equiv \langle 0,0|v|1,1 \rangle,$$

(23)

with $|0,0\rangle$ and $|1,1 \rangle$ the spherical harmonic functions $Y_{0,0}$ and $Y_{1,1}$, respectively. $a_{cv}$ has the dimension of length. $U$, $K$, and $Q$ in Eqs. (21) and (22) have been defined in Eqs. (7) and (10). Note that the optical selection rules owing to the $s$- and $p$-wave natures of the conduction and valence bands have been incorporated in Eqs. (21) and (22). We neglect the density matrix of the valence bands because the charge and spin lifetimes of holes are much shorter than those of electrons for $n$-type quantum wells.

In this work we will retain only the diagonal part of the density matrix because when ignoring the broadening of the light frequency $\omega$, the coherent contribution from the off-diagonal part of the density matrix can be neglected. The diagonal terms of the density matrix have clear physical meaning as the photoexcited carrier density. By substituting the polarization electric field vector $\mathbf{E} = E_0(\cos \theta, \sin \theta)$ into Eq. (20), the carrier density excited to conduction band $|c,k\rangle$ can be written as the summation from different valence bands $|v,k\rangle$:

$$\rho_{c,c}(k) = \sum_v \rho_{c,v,k},$$

(24)

where $\rho_{c,v,k}$ represents the carrier density excited from valence band $|v,k\rangle$ to conduction band $|c,k\rangle$ as a function of the wavevector $k$, $\rho_{c,v,k}$ can be divided into three terms according to their dependence on the polarization angle $\theta$,

$$\rho_{c,v,k} = \rho_{c,v,k}^{(0)} + \rho_{c,v,k}^{\cos} \cos 2 \theta + \rho_{c,v,k}^{\sin} \sin 2 \theta,$$

(25)

where

$$\begin{align*}
\rho_{c,v,k}^{(0)} &= \xi I \frac{2 \pi \tau_c \epsilon_c^2}{\hbar^2} \sqrt{\frac{\mu_{0}}{\epsilon_0}} (1 - f_c)\delta(\omega - \omega_{c,v})(|x_{c,v}|^2 + |y_{c,v}|^2), \\
\rho_{c,v,k}^{\cos} &= \xi I \frac{2 \pi \tau_c \epsilon_c^2}{\hbar^2} \sqrt{\frac{\mu_{0}}{\epsilon_0}} (1 - f_c)\delta(\omega - \omega_{c,v})(|x_{c,v}|^2 - |y_{c,v}|^2), \\
\rho_{c,v,k}^{\sin} &= \xi I \frac{2 \pi \tau_c \epsilon_c^2}{\hbar^2} \sqrt{\frac{\mu_{0}}{\epsilon_0}} (1 - f_c)\delta(\omega - \omega_{c,v})\text{Re}(2x_{c,v}y_{c,v}).
\end{align*}$$

(26)
direct interband transitions from valence band and valence bands also have similar anisotropic photoexcited carrier density from field along the bands have different carrier density on the spin-split conduction band. The states involved in the transitions form a set of constant energy contours (denoted by dashed rings), which can be shifted along \( k_x \) in the presence of a magnetic field along the \((x,y)\) direction.

In the polar coordinates \((k_x, k_y) \equiv k(\cos \varphi, \sin \varphi)\), we can transform the \( \delta \) function of \( \omega \) into that of \( k \):

\[
\rho_{cv, k}^0 = \xi I \frac{2\pi \tau_e \varepsilon_T}{\hbar^2} \frac{\mu_0}{\varepsilon_0} (1 - f_c) G[k_{cv}(\varphi)] \delta[k - k_{cv}(\varphi)] \\
\times (|x_{cv}|^2 + |y_{cv}|^2), \\
\rho_{cv, k}^{\cos} = \xi I \frac{2\pi \tau_e \varepsilon_T}{\hbar^2} \frac{\mu_0}{\varepsilon_0} (1 - f_c) G[k_{cv}(\varphi)] \delta[k - k_{cv}(\varphi)] \\
\times (|x_{cv}|^2 - |y_{cv}|^2), \\
\rho_{cv, k}^{\sin} = \xi I \frac{2\pi \tau_e \varepsilon_T}{\hbar^2} \frac{\mu_0}{\varepsilon_0} (1 - f_c) G[k_{cv}(\varphi)] \delta[k - k_{cv}(\varphi)] \\
\times \Re(2x_{cv}y_{cv}),
\]

where we defined

\[
G[k_{cv}(\varphi)] \equiv \frac{1}{|d\omega_{cv}(k,\varphi)/dk|} |_{k=k_{cv}(\varphi)}.
\]

**FIG. 2.** (Color online) Schematic description of direct interband optical transitions from a valence band \(|v, k\rangle\) to the spin-split conduction bands \(|c, \pm k\rangle\) at zero magnetic field (a) and in the presence of a magnetic field \(B\) (b). The states involved in the transitions form a set of constant energy contours (denoted by dashed rings), which can be shifted along \( k_x \) in the presence of a magnetic field.

**FIG. 3.** (Color online) At zero magnetic field \((B_x = 0, B_y = 0)\), the calculated photoexcited carrier density \(\rho_{cv, k}\) in Eq. (27) as functions of the wavevector angle \(\varphi\). Dark (light) represents positive (negative) values. “max” (“min”) indicates the maximum (minimum) values. The \(\rho_{cv, k}^0\) term is always positive and overwhelms \(\rho_{cv, k}^{\cos, \sin}\). \(\rho_{cv, k}^{\cos, \sin}\) are in units of \((2\pi \xi I \tau_e \varepsilon_T \mu_0/\hbar) \times \mathrm{second/meter}\), and have a dimension of meter\(^{-1}\). Parameters: \(T = 77\ \mathrm{K}, E_\gamma = 0.01\ \mathrm{eV}, \omega = 0.8\ \mathrm{eV}, m^* = 0.04 m_0, \gamma_1 = 11.97, \gamma_2 = 4.36, \Delta = 0.764\ \mathrm{eV}, d = 40 \ \mathrm{nm}, \alpha = 4.31\ \mathrm{meV} \cdot \mathrm{nm}, \beta = 0\). Numerical scheme is given in Appendix B.

In this case, Eq. (29) can be written as

\[
\rho_{cv, k} = \rho_{cv, k}^0 + \rho_{cv, k}^{\cos} \cos 2\varphi \cos 2\theta + \rho_{cv, k}^{\sin} \sin 2\varphi \sin 2\theta,
\]

where \(\rho_{cv, k}^{\cos, \sin}\) in general are functions of \(\varphi\). In a weak in-plane magnetic field (say, less than 1 tesla), \(\rho_{cv, k}^{\cos, \sin}\) slightly depend on \(\varphi\). In the absence of the magnetic field and the anisotropy of the spin-orbit couplings, \(\rho_{cv, k}^{\cos, \sin}\) become independent on \(\varphi\). In this case, Eq. (29) can be written as

\[
\rho_{cv, k} = \rho_{cv, k}^0 + \rho_{cv, k}^{\cos} \cos 2\varphi \cos 2\theta + \rho_{cv, k}^{\sin} \sin 2\varphi \sin 2\theta,
\]

where \(\rho_{cv, k}^{\cos, \sin}\) are constants of dimension meter\(^{-1}\).

The anisotropy of the photoexcited carrier density is the core of this paper. It can naturally account for the field and polarization dependence of the magnetoelectric photocurrent\(^{25}\) and the pure spin photocurrent.\(^{20,21}\)

**V. MAGNETIC-FIELD INDUCED ELECTRIC PHOTOCURRENT**

**A. Origin of \(c_v, c_x,\) and \(c_y\)**

The electric photocurrent density along \(\mu \in \{x,y\}\) axis can be found by summing the velocities of the photoexcited carriers

\[
\vec{j}_\mu = -e \sum_{c_v, k} \rho_{cv, k} \vec{v}_{cv, k}^\mu.
\]
From the eigenenergies Eq. (6), the velocity along the $x$ direction for conduction bands $|\pm, k\rangle$ can be found as

$$v_{x,k}^\pm = \frac{1}{h} \frac{\partial E_k}{\partial k_x} = \frac{\hbar}{m^*} k_x \pm \frac{\alpha}{\hbar} (\alpha k_x - h_y) \sqrt{(\alpha k_x + h_y)^2 + (\alpha k_x - h_y)^2}.$$  

(32)

Above and hereafter, we will ignore $\beta$ in the analytical expressions, and take it into account only in the numerical calculation. By rewriting $k_x = k \cos \varphi$ and $k_y = k \sin \varphi$ in polar coordinates $(k, \varphi)$ and expanding the velocities up to the linear order in $h_x$ and $h_y$, we have

$$v_{x,k}^\mu = \left( \frac{\hbar}{m^*} k \pm \frac{\alpha}{\hbar} \right) \cos \varphi \mp \frac{\sin 2\varphi}{2k} h_x \mp \frac{\sin^2 \varphi}{kh} h_y.$$  

(33)

$v_{x,k}^\mu$ in the current density formula (31) is related to Eq. (33) by restricting $k$ on the constant energy contours $k_{cv}$. Rewriting the summation in Eq. (31) into an integral in polar coordinates, putting into Eqs. (29) and (33), and performing the integral over $k$, we obtain an integral over $\varphi$:

$$j_x = -\frac{e}{(2\pi)^2} \sum_{c,v} \int^{2\pi}_0 d\varphi k_{cv}(\varphi) \times \left( \rho^0_{c,v} + \rho^\cos_{c,v} \cos 2\varphi \cos 2\theta + \rho^\sin_{c,v} \sin 2\varphi \sin 2\theta \right) $$

$$\times \left[ \left( \frac{\hbar}{m^*} k_{cv}(\varphi) + \frac{\alpha}{\hbar} \right) \cos \varphi - \frac{c}{2hk_{cv}(\varphi)} h_x - \frac{\sin 2\varphi}{khk(\varphi)} h_y \right].$$  

(34)

where $\psi \equiv \arctan(k_x/k_z)$, $k^2 \equiv k_x^2 + k_y^2$, and $k_{cv}(\varphi)$ are the roots of $\omega = \omega_{cv}(k, \varphi)$. In general, $k_{cv}(\varphi)$ and $\rho_{c,v}^{0,\text{cos, sin}}$ are functions of $\varphi$ when $B \neq 0$ and $\beta \neq 0$. At this moment, if we ignore their $\varphi$ dependence, the above integral can be readily performed:

$$j_x(B_x, B_y) \simeq h_y \frac{e}{(2\pi)^2} \left( \sum_{c,v} c \rho^0_{c,v} \right) \int^{2\pi}_0 d\varphi \sin^2 \varphi + h_x \frac{e}{(2\pi)^2} \left( \sum_{c,v} c \rho^\cos_{c,v} \right) \cos 2\theta \int^{2\pi}_0 d\varphi \cos 2\varphi $$

$$+ \frac{e}{4\pi h} \left( \sum_{c,v} c^0_{cv} \right) h_y + \frac{e}{8\pi h} \left( \sum_{c,v} c_{cv} \right) h_x \sin 2\theta.$$  

(35)

The resulting three nonzero terms immediately recover the form of the experimental fitting formula (1). In this way, a clear relation between the anisotropy of the photoexcited carrier density and the magnetoelectric photocurrent density is established.

Above we ignore the $\varphi$ dependence of $\rho_{c,v}^{0,\text{cos, sin}}$ and $k_{cv}$ resulting from finite $B$ and $\beta$. Their influence can be taken into account numerically. In Fig. 4 we present a product table of calculated $v_{x,k}^\mu$ and $\rho_{c,v}^{0,\text{cos, sin}}$ as functions of $\varphi$, for finite $B$ and $\beta$ (refer to Appendix B for the numerical scheme). The three nonzero contributions in Eq. (35) are marked by the boxed subfigures (iic), (iia), and (iiib). Except the three nonzero terms, note that all the other subfigures in Fig. 4 always have equal weight of positive and negative lobes. In other words, they yield zero when integrating over $\varphi$, and give no contribution to the electric current. All the pairs of conduction and valence bands have the same behavior, and all of them add up (actually cancel with each other) to give the total net electric current.

In the analytic result (35), we have neglected the variation of $\rho_{c,v}^{0,\text{cos, sin}}$ when applying the magnetic field. In general, the charge current $j_{\mu}$ induced by $B_{\nu}$ can be expanded as

$$j_{\mu}(B_{\nu}) \simeq -e \sum_{c,v,k} \left[ \rho_{c,v,k}(0) v_{x,k}^\mu(0) + \rho_{c,v,k}(0) \Delta v_{x,k}^\mu(B_{\nu}) \right]$$

$$+ \Delta \rho_{c,v,k}(B_{\nu}) v_{x,k}^\mu(0) + \Delta \rho_{c,v,k}(B_{\nu}) \Delta v_{x,k}^\mu(B_{\nu}) \]$$

$$+ O(B_{\nu}^2),$$  

(36)

where “(0)” stands for “at zero magnetic field”, and “$\Delta \ldots(B_{\nu})$” for the variation when applying the magnetic field $B_{\nu}$ along the $\nu$ direction. We already illustrated that the first term vanishes, and the second term is consistent with the experiment. Besides, one can expect that the third term, i.e., the magnetic field-induced variation of the photoexcited carrier density, gives a contribution of the same order as the second term (see Fig. 5), and the last term is ignorably small. Therefore, we approximate the photocurrent density induced by the magnetic field by

$$j_{\mu}(B_{\nu})$$

$$\simeq \left( \eta \frac{e}{2\pi h} \sum_{c,v} c_{cv} \right) h_y - \left( \eta \frac{e}{4\pi h} \sum_{c,v} c_{cv} \right) h_x \cos 2\varphi $$

$$+ \left( \eta \frac{e}{4\pi h} \sum_{c,v} c_{cv} \right) h_x \sin 2\varphi,$$  

(37)

where an extra “$2\eta$” has been multiplied to account for the summation of the second and third terms of Eq. (36).

B. Comparison with experiment

Above we show that the photoexcited carrier density in Eq. (29) explains the origin of $c_0$, $c_\gamma$, and $c_\varepsilon$ terms in the fitting formula (1) of the magnetoelectric photocurrent. Now we compute $c_0/\sqrt{\gamma}$, and see how close the minimal model can be when comparing with the experiment.
We choose a set of parameters close to the experiment. From the experiment, we have that the temperature $T = 77 \text{ K}$, the Fermi level $E_F = 0.01 \text{ eV}$, the light frequency $\omega = 0.8 \text{ eV}$, the band gap $\Delta \approx 0.764 \text{ eV}$, the quantum well thickness $d = 40 \text{ nm}$, and the Rashba spin-orbit coupling constant $\alpha = 4.31 \text{ meV} \cdot \text{nm}$. The Luttinger model parameters $\gamma_1 = 11.97$ and $\gamma_2 = 4.36$ are adopted from those for GaAs, which has the similar components as those in the experiment (Ga$_{1-x}$In$_x$As, $x = 0.53 \sim 0.59$ by graded doping). The momentum relaxation time can be estimated by $\tau_e = \frac{m^* \nu_e}{\epsilon} \approx 2 \text{ ps}$, with the effective mass $m^* = 0.04$ and the mobility $\mu_e \approx 84,000 \text{ cm}^2/\text{V} \cdot \text{s}$ at 77 K (about 7 times larger than that at room temperature). The effective dipole length $a_v \approx 6.7 \text{ Å}$ is approximated by that for GaAs. Considering a reflectance of 0.3 and the absorption coefficients $9 \times 10^3 \text{ cm}^{-1}$, we obtained the absorption efficiency $\xi = (1 - 0.3)[1 - \exp(-9 \times 10^3 \text{ cm}^{-1} \times 40 \text{ nm})] \approx 2.5\%$. With the light power $P = 15 \text{ mW}$ and the light spot radius $5 \mu\text{m}$, the light intensity is found as $I \approx 1.91 \times 10^9 \text{ W} / \text{m}^2$. The experiment observed that $\xi \neq |\xi|$; this may be due to a finite Dresselhaus spin-orbit coefficient $\beta$, so that the spin-orbit couplings along the $x$ and $y$ directions are different.
i.e., $\lambda_x \neq \lambda_y$, since $\lambda_x = \alpha - \beta$ and $\lambda_y = \alpha + \beta$. The value of $\beta$ is unknown to the experiment. In this following calculations, $\beta$ will be chosen as a variable parameter.

In Fig. 6, we present the numerically calculated photocurrent densities $j_x$ as functions of $B_x$, $B_y$, and $\theta$. We also compare the results for positive, zero, and negative $\beta$. Qualitatively, the
calculated results capture the main features of the experiment. Figs. 6(b) and 6(d) show that the current densities are linearly scaled with the magnetic fields. At zero magnetic field, there is no current. Figures 6(a) and 6(c) show that the parallel field leads to only sin $2\theta$ dependence, while the perpendicular field induces only the constant and cos $2\theta$ dependence. In summary, the calculated current density can be formulated in the same form as Eq. (1).

Now we make some quantitative comparisons. In Fig. 6, the current densities are in units of $\zeta \approx 0.09$ A/m [see Eq. (B2)]. Figure 6 shows that the dominant term is $j_\theta(B_x)$, consistent with the experiment. $j_\theta(B_x)$ can be as large as $8 \times 10^{-5} \zeta$ when $B_x = 1$ T, i.e., about $0.72 \times 10^{-5}$ A/m, comparable with the experimental estimate $\sim 2 \times 10^{-5}$ A/m.$^{25}$ The calculated current is smaller, probably because only limited bands are included in the calculation. In Table I, we compare the calculated $c_0$, $c_y$, and $c_z$ with the experiment. The calculated $c_y$ and $c_z$ are comparable with $c_0$, while in the experiment $c_0$ is almost an order of magnitude larger than $c_y$ and $c_z$. In Sec. V we have shown that $c_y$ and $c_z$ terms come from the photoexcited carrier $\rho_{\text{spin}}^{\text{cos}, \sin}$, which originate from the quantum interference between two circular components of a linearly polarized light.$^{16}$ In contrast to the ideal situation assumed in the theory, the interference effect may be suppressed in the experiment; then $c_y$ and $c_z$ are reduced and $c_0$ is enhanced by a noninterference contribution. Besides, the sign of the calculated $c_y$ is opposite to that of the experiment. This may be attributed to the difference in band symmetries between the experiment and the model. Considering the simplicity of the minimal model, we expect that more sophisticated models may cover more reliable details, but the minimal model is enough to offer a reasonable physical picture for the experiment.

VI. ZERO-FIELD PURE SPIN PHOTOCURRENT

A. Polarization dependence of zero-field pure spin photocurrent

Both the symmetry argument of the $C_{2v}$ group$^{17}$ and theoretical calculations$^{16,20,21}$ have pointed out the generation of pure spin currents by the linearly polarized or unpolarized lights, and they can be expressed as functions of $\theta$:

$$j_{\mu}^y = I_0 + I_1 \cos 2\theta, \quad j_{\mu}^z = I_2 \sin 2\theta,$$

where $I_{0,1,2}$ are constant coefficients. In this section, we will show how the anisotropy of the photoexcited carrier density is related to the $\theta$ dependence of the pure spin current.

The spin photocurrent density flowing along the $\mu (\in \{x, y\})$ direction with spin polarized along the $v (\in \{x, y\})$ direction can be found by

$$j_{\mu v} = \frac{\hbar}{2} \sum_{c, v, k} \rho_{c v, k} j_{\nu v, k}^{(c)},$$

where $j_{\nu v, k}^{(c)}$ are coefficients that depend on $c, v, k$. In Sec. V B.
where we also retain only the diagonal density matrix, and \( s^{\mu\nu}_{c,-k} \) is the spin velocity for the carriers excited from valence band \([v,k]\) to conduction band \([c,k]\). The spin velocity operator is defined as
\[
\hat{\mathbf{v}}^{\mu\nu} = \frac{1}{2} \left\{ \sigma_\mu, \frac{1}{\hbar} \partial \mathbf{H}_C \right\},
\]
(40)
where \( \sigma_\mu \) is the Pauli matrix. The zero magnetic field expectation value of the spin velocity for conduction band \([c,k]\), defined as \( s^{\mu\nu}_{c,k} = \langle \pm, k | \hat{\mathbf{v}}^{\mu\nu} | \pm, k \rangle \), can be found in polar coordinates as
\[
s^{x^x}_{\pm k} = \mp \frac{\hbar}{2m^*} k \sin 2\varphi, \quad s^{xy}_{\pm k} = \mp \frac{\hbar}{2m^*} k \cos^2 \varphi - \frac{\alpha}{\hbar},
\]
(41)
\[
s^{y^y}_{\pm k} = \mp \frac{\hbar}{2m^*} k \sin^2 \varphi + \frac{\alpha}{\hbar}, \quad s^{yx}_{\pm k} = \mp \frac{\hbar}{2m^*} k \sin 2\varphi.
\]

\( s^{\mu\nu}_{c,-k} \) in the spin current density formula (39) is related to Eq. (41) by restricting \( k \) on the constant energy contours between valence band \([v,k]\) and conduction band \([c,k]\).

Similar to the current density, the spin current density can also be rewritten into an angle integral,
\[
j^{y^y}_{c,\rho} (0) = \frac{\hbar}{2(2\pi)^2} \sum_{c,v} \int_0^{2\pi} d\varphi k_{cv}(\varphi) s^{xy}_{c,\rho} (k_{cv}, \varphi)
\]
\[
\times \left( \rho_{cv,0}^0 + \rho_{cv,0}^\cos \cos 2\varphi \cos 2\theta + \rho_{cv,0}^\sin \sin 2\varphi \sin 2\theta \right),
\]
(42)
If we ignore the \( \varphi \) dependence of \( \rho_{cv,0}^0, \cos, \sin \) and \( k_{cv}(\varphi) \), the above integral yields
\[
j^{y^y}_{c,\rho} (0) = \frac{\hbar}{2(2\pi)^2} \sin 2\theta \sum_{c,v} \frac{\hbar}{2m^*} k^2_{cv} \rho_{cv}^\sin,
\]
\[
j^{y^y}_{c,\rho} (0) = -\frac{\hbar}{2(2\pi)^2} \sum_{c,v} \left( \frac{\hbar}{2m^*} k^2_{cv} \rho_{cv}^0 + \frac{\alpha}{\hbar} \rho_{cv}^0 \right)
\]
\[
-\frac{\hbar}{2(2\pi)^2} \sum_{c,v} \frac{\hbar}{2m^*} k^2_{cv} \rho_{cv}^\cos \cos 2\theta,
\]
(43)
which give the \( \theta \) dependence in Eq. (38). In the presence of a finite \( \beta \), the result is not affected qualitatively, as shown in Fig. 7, where the nonzero contributions to Eq. (38) are marked by the boxes.

**B. Quick estimate of zero-field spin photocurrent from magnetoelectric photocurrent**

Because both the magnetoelectric photocurrents and the zero-field pure spin photocurrents originate from the same photoexcited carrier density, this allows us to find a relation between them. With the help of Eqs. (37) and (43), the ratio of the zero-field longitudinal pure spin photocurrent to the electric photocurrent induced by the parallel magnetic field turns out to be
\[
\left| \frac{j^{y^y}_{c,\rho} (0)}{j_{c}(B_f)} \right| \approx \frac{\sum_{c,v} \rho_{cv}^\sin k^2_{cv}}{\eta h \sum_{c,v} \rho_{cv}^\sin},
\]
(44)
and the ratio of the transverse spin photocurrent to the perpendicular magnetic field induced electric photocurrent is given by
\[
\left| \frac{j^{y^y}_{c,\rho} (0)}{j_{c}(B_f)} \right| \approx \frac{\sum_{c,v} \rho_{cv}^\sin k^2_{cv}}{\eta h \sum_{c,v} \rho_{cv}^\sin} \left[ \frac{\alpha^2}{2m^*} + \frac{\hbar^2 k^2_{cv}}{2m^*} \right]
\]
(45)
These relations have clear physical meaning. Note that \( \sum_{c,v} \rho_{cv,0}^\cos, \sin \rho_{cv,0}^\cos, \sin \rho_{cv,0}^\cos, \sin \) . This difference of the photocarrier density between the + and − conduction bands, as already shown in Fig. 3, is due to the spin-orbit coupling. Therefore, the denominators in Eqs. (44) and (45) mean that both the magnetic field and the spin-orbit coupling are necessary ingredients of the magnetoelectric photocurrent, while the numerators indicate that the pure spin currents are proportional to the spin-orbit coupling and the kinetic energy of the photoexcited carriers (if we can view literally \( \hbar^2 k^2_{cv} / 2m^* \) as kinetic energy).

At this moment, we make a bold approximation by canceling the effect of the spin-orbit coupling from both the denominators and numerators in Eqs. (44) and (45), and literally say that the ratio of the zero-field pure spin photocurrent to the magnetoelectric photocurrent is about "kinetic energy over Zeeman energy." This relation, though rather coarse, can help us to make a quick order-of-magnitude estimate of the undetectable pure spin photocurrent from the measured magnetoelectric photocurrent. The average kinetic energy of the photoexcited carriers is higher than the equilibrium Fermi energy measured from the bottom of the conduction bands, and thus is more than \( E_F = 0.01 \) eV for our numerical calculations. The Zeeman energy induced by 1 tesla of magnetic field is about \( 10^{-4} \) eV for the Landé g factor of \( g_e = -4.34-36 \). Therefore, the rough estimate implies that the spin photocurrent is about two orders larger than the magnetoelectric photocurrent at 1 tesla. To test the reliability of this quick estimate, we numerically compare the magnetoelectric photocurrent and zero-field pure spin photocurrent in Fig. 8. The spin photocurrents are about 250 times larger than the magnetoelectric photocurrent at 1 tesla of magnetic field. Despite its roughness, this quick estimate gives a reasonable result. We expect that this quick estimate can serve as a reference for more sophisticated nondestructive approaches, such as the Faraday rotation or the second-order nonlinear optical effects.

**VII. CONCLUSIONS**

In this work we present a theoretical description of the recent experiment on the optical injection of spin-polarized carriers via direct interband optical excitations into a semiconductor quantum well under the normal incidence of linearly polarized or unpolarized lights. In that experiment the injection produces pure spin photocurrents accompanying no electric current at zero magnetic field due to spin-orbit coupling that respects time-reversal symmetry. An in-plane magnetic field can break time-reversal symmetry and extracts a magnetoelectric photocurrent. The magnetoelectric photocurrent is
characterized by its dependence on the magnetic fields and the polarization of the light.

With the help of the density matrix formalism, we calculate the photoexcited carrier density, current density, and spin current density. The photoexcited carrier density in $k$ space shows an anisotropic dependence on both the wavevector angle $\varphi$ and the polarization angle $\theta$ of the linearly polarized light as given in Eq. (30). Since the velocities of carriers can also be expressed as functions of $\varphi$ [see Eq. (33)], we can show that the current density can be simplified as an angle integral over $\varphi$ of the product of the density and velocity of photoexcited carriers. The angle integral then produces all the magnetic field and polarization angle dependence of the magnetoelectric photocurrent as reported in Ref. 25, in particular, the magnetoelectric photocurrent induced by the parallel magnetic field.

We show that the present simplified model with Rashba and Luttinger Hamiltonians is able to reproduce the current formula for the magnetoelectric photocurrent. However, discrepancies still exist between theory and experiment for the relative magnitudes and signs among the parameters in the current formula. Since the simplified model is a natural choice considering the symmetry of the quantum well that was investigated in the experiment, the discrepancies indicate that further investigations, with more detailed considerations on the band structure of the sample, are needed to identify the origin of the discrepancies. For example, more energy bands may be required in the calculation, the graded doping in the quantum well may need a self-consistent calculation of the potential and carrier density.

We further show that the origin of the previously predicted pure spin photocurrents\(^{20,21}\) can be well illustrated from the same photoexcited carrier density. We propose that the ratio of the zero-field pure spin photocurrent to the magnetoelectric photocurrent can be approximated as “kinetic energy over Zeeman energy.” With this relation, the underlying pure spin photocurrent can be quickly estimated from the observed magnetoelectric photocurrent and provides a reference for other approaches to nondestructive measurement of the pure spin photocurrent.

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reflection planes of the are defined as \([1\overline{1}0]\) and \([110]\) crystallographic directions, respectively, the \(xz\) nonlinear optics. Phenomenologically, the current density representation \(B_1\) of the field) along the \(x\) direction have the asymmetric zincblende materials grown along \([001]\) crystallographic directions. When \(x\) and \(y\) axes are defined as \([110]\) and \([100]\) crystallographic directions, the \(xz\) and \(yz\) planes coincide with the mirror reflection planes of the \(C_{2v}\) group. Two examples are illustrated in Table II. In the language of the irreducible representation of group theory, Table II can be written as

\[
B_1 \neq B_2 \otimes B_1 \otimes B_1,
\]

Similarly, all the nonzero terms can be found and summarized as Eq. (2).

In addition, both \(\sigma_x\) and \(k_x\) transform according to \(B_1\), while \(B_1 \otimes B_1\) yields the identity representation of the \(C_{2v}\) group, so \(\lambda_x k_x\) is an invariant for structures with the \(C_{2v}\) symmetry. Similarly, \(\sigma_x k_x\) is also an invariant. On the contrary, \(\sigma_y k_x\) and \(\sigma_y k_x\) are not invariants. As a result, the spin-orbit coupling up to the linear order in \(k\) can be generally described by the form

\[
H_{\text{SOC}} = \lambda_x \sigma_y k_x - \lambda_x \sigma_y k_x,
\]

where \(\lambda_x\) and \(\lambda_x\) are spin-orbit coupling coefficients along different directions. In the coordinate system where \(x'[[100]\) and \(y'[[010]\), the spin-orbit coupling in the conduction bands of a sample with the \(C_{2v}\) symmetry can be generally written...
TABLE II. Two examples of how to determine whether an element of the pseudotensor $\chi^{\alpha\beta\gamma}$ is nonzero. 1 and $-1$ correspond to “remaining unchanged” and “changing sign,” respectively, upon applying the symmetry operations of the $C_{\nu}$ group to the vectors. (a) For $\chi^{x\alpha\alpha}$, which is zero because $j_{x}$ and $B_{x}E_{y}$ are different for $\sigma_{x}$ and $\sigma_{y}$. (b) For $\chi^{x\alpha\nu}$, which is nonzero because $j_{x}$ and $B_{x}E_{y}$ are the same for all symmetry operations.

<table>
<thead>
<tr>
<th>$j_{x}$</th>
<th>$B_{x}$</th>
<th>$E_{x}$</th>
<th>$E_{y}$</th>
<th>$B_{x}E_{y}$</th>
</tr>
</thead>
<tbody>
<tr>
<td>$\sigma_{x}$</td>
<td>$-1$</td>
<td>$-1$</td>
<td>$-1$</td>
<td>$-1$</td>
</tr>
<tr>
<td>$\sigma_{y}$</td>
<td>$1$</td>
<td>$-1$</td>
<td>$1$</td>
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<tr>
<td>$\sigma_{x}$</td>
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<td>$-1$</td>
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<tr>
<td>$\sigma_{y}$</td>
<td>$1$</td>
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</tr>
</tbody>
</table>

as

$$H_{SOC} = \alpha(\sigma_{x}k_{y} - \sigma_{y}k_{x}) + \beta(\sigma_{x}k_{y} - \sigma_{y}k_{x}),$$

(A4)

referred to as Rashba ($\alpha$) and Dresselhaus ($\beta$) terms, respectively. While in this work the coordinate system is $x | [110]$ and $y | [110]$, the form of the spin-orbit coupling can be obtained by rotating the above $H_{SOC}$ by $45^\circ$:

$$H_{SOC} = (\alpha + \beta)\sigma_{x}k_{y} - (\alpha - \beta)\sigma_{y}k_{x}.$$  

(A5)

The extra $\beta$ means that although the $C_{\nu}$ symmetry allows only $\sigma_{x}k_{x}$- or $\sigma_{y}k_{y}$-type spin-orbit coupling when $x | [110]$ and $y | [110]$, the Dresselhaus term when $x | [100]$ and $y | [010]$ can lead to an anisotropy of the spin-orbit coupling. Usually, $\beta$ is smaller than $\alpha$.

Note that this result applies only for when $x$ and $y$ directions are referred to the [110] and [110] crystallographic directions.

APPENDIX B: NUMERICAL SCHEME

For each pair of conduction ($c$) and valence ($v$) bands, we denote all the quantum states by the discrete values of the wavevector angle $\varphi = \varphi_{i} = i2\pi/N$, with $i = 0,1, \ldots, N$. At each $\varphi_{i}$, the energy conservation law $\omega = \omega_{c}(\varphi_{i})$ is numerically solved. The root of $k$ is the wavevectors $k_{\nu}(\varphi_{i})$ on the constant energy contours. The physical quantities such as $v_{c}^{\nu}(k_{\nu}(\varphi_{i}))$ and $\rho_{cv}^{\nu}(\varphi_{i})$ are calculated by giving them $k = k_{\nu}(\varphi_{i}) (\cos \varphi_{i}, \sin \varphi_{i})$, and denoted as $v_{c}^{\nu}(k_{\nu}(\varphi_{i}))$ and $\rho_{cv}^{\nu}(\varphi_{i})$. As an example, we show how to numerically calculate the $\theta$-independent term of $j_{x}$. According to Eqs. (34) and (27), it means that we calculate

$$j_{x}^{0} = -\frac{\xi I_{e} e_{a}e_{c}^{2}\sqrt{\mu_{0}/\varepsilon_{0}}}{2\pi \hbar^{2}} \sum_{c,v}^{N-1} \Delta k_{cv}(\varphi_{i}) e_{cv}^{x} \left(\frac{|x_{cv}|^{2} + |y_{cv}|^{2}}{a_{cv}^{2}}\right)_{k_{\nu}(\varphi_{i})} G[k_{\nu}(\varphi_{i})],$$

(B1)

where $\Delta k_{cv}(\varphi_{i}) = \Delta_{v}(\varphi_{i}) - \Delta_{c}(\varphi_{i}) - 1$, and $\sum_{c,v}^{N-1} \Delta k_{cv}(\varphi_{i})$ gives the circumference of the constant energy contour. In the summation, we are using the units $[G] \sim$ second/meter, $[v^{\nu}] \sim$ meter/second, $[\Delta k_{cv}] \sim$ nm$^{-1}$, so the units of the current density in front of the dimensionless summation of Eq. (B1) are

$$[j_{x}^{0}] \sim \frac{\xi I_{e} e_{a}e_{c}^{2}\sqrt{\mu_{0}/\varepsilon_{0}}}{2\pi \hbar^{2}} \approx 0.09$$.  

(B2)
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